# Analysis of the motion of a non-holonomic mechanical system 

## D.N. Zeković

Beograd, Serbia

## A R T I C L E I N F O

## Article history:

Received 25 April 2006


#### Abstract

The motion of a plane non-holonomic mechanical system, consisting of two point masses, which move in such a way that their velocities are mutually perpendicular, is analysed [Zeković D. Examples of non-linear non-holonomic constraints in classical mechanics. Vestnik MGU. Ser. 1. Matematika Mekhanika, 1991; $1: 100-3]$. The equations of the constraints of such a system are derived, the reactions of the constraints are calculated and the cyclical first integrals are written.


© 2008 Elsevier Ltd. All rights reserved.

## 1. Analysis of the constraint equations

The practical realization of a system consisting of two point masses, which move in such a way that their velocities are mutually perpendicular, can be achieved using two knife edges and a weightless rigid construction ("a fork"), as shown in the Fig. 1. This system is non-holonomic ${ }^{1}$.

Suppose $x_{1}, y_{1}, x_{2}$ and $y_{2}$ are generalized coordinates. The condition for the velocities of the point masses to be orthogonal can be written in the form

$$
\begin{equation*}
\dot{x}_{1} \dot{x}_{2}+\dot{y}_{1} \dot{y}_{2}=0 \tag{1.1}
\end{equation*}
$$

Two other constraint equations express the conditions for the velocities of the points $M_{1}$ and $M_{2}$ and the planes of the knife edges to be codirectional:

$$
\begin{equation*}
\dot{x}_{1} \cos \varphi+\dot{y}_{1} \sin \varphi=0, \quad \dot{x}_{1} \sin \varphi-\dot{y}_{2} \cos \varphi=0 \tag{1.2}
\end{equation*}
$$

Constraint equations (1.1) and (1.2) are dependent, and Eq. (1.1) and any of Eqs. (1.2) or both of Eqs. (1.2) can be used. Since the reactions of the constraints $\mathbf{R}_{1}$ and $\mathbf{R}_{2}$ are perpendicular to the planes of the knife edges, their scalar product is equal to zero.

The left-hand side of Eq. (1.1) is a homogeneous function of $\Phi$ of power two, and this relation can be written in the form

$$
\frac{\partial \Phi}{\partial \dot{q}^{i}} \dot{q}^{i}=0
$$

where $q^{i}$ are generalized coordinates, and the possible displacements permitted by this non-linear constraint, we will define, like Chetayev, in the form

$$
\frac{\partial \Phi}{\partial \dot{q}^{i}} \delta \dot{q}^{i}=0
$$

Here and henceforth summation is carried out over repeated indices, where

$$
i, j, k=1, \ldots, n ; \quad \alpha, \beta, \gamma=1, \ldots, m ; \quad \nu, \rho=m+1, \ldots, m+l=n
$$

Note that the coefficients of $\delta q^{i}$ depend on the velocities.

[^0]0021-8928/\$ - see front matter © 2008 Elsevier Ltd. All rights reserved. doi:10.1016/j.jappmathmech.2008.11.012


Fig. 1.

## 2. The trajectories of points of the system

We will write the equations of motion of the system, choosing as the generalised coordinates $x_{\mathrm{i}}, y_{\mathrm{i}}, \xi$ and $\varphi$ (see the Fig. 1),

$$
\begin{equation*}
\dot{x}_{1} \cos \varphi+\dot{y}_{1} \sin \varphi=0, \quad \dot{x}_{1} \sin \varphi-\dot{y}_{1} \cos \varphi-\xi \dot{\varphi}=\sqrt{\dot{x}_{1}^{2}+\dot{y}_{1}^{2}}-\xi \dot{\varphi}=0 \tag{2.1}
\end{equation*}
$$

Taking $\dot{\xi}$ and $\dot{\varphi}$ as independent velocities, we obtain

$$
\begin{equation*}
\dot{x}_{1}=\xi \dot{\varphi} \sin \varphi, \quad \dot{y}_{1}=-\xi \dot{\varphi} \cos \varphi \tag{2.2}
\end{equation*}
$$

The kinetic energy of the system, provided that the masses of the points $M_{1}$ and $M_{2}$ are equal to unity, can be written in the form

$$
\begin{equation*}
T=\frac{1}{2}\left(\dot{x}_{1}^{2}+\dot{y}_{1}^{2}\right)+\frac{1}{2}\left(\dot{x}_{2}^{2}+\dot{y}_{2}^{2}\right), \quad T^{*}=\frac{1}{2}(\xi \dot{\varphi})^{2}+\frac{1}{2}(\dot{\xi})^{2} \tag{2.3}
\end{equation*}
$$

where $T^{*}$ is the result of eliminating the dependent velocities in the expression for $T$ using the constraint equations.
The system is a Chaplygin system in its inertial motion, and its equations of motion in Chaplygin form are ${ }^{2,3}$

$$
\begin{align*}
& \frac{d}{d t} \frac{\partial T^{*}}{\partial \dot{q}^{\alpha}}-\frac{\partial T^{*}}{\partial q^{\alpha}}+\frac{\partial T^{*}}{\partial \dot{q}^{v}} \gamma_{\alpha}^{v}=0 \\
& \gamma_{\alpha}^{v}=\frac{\partial \psi^{v}}{\partial q^{\alpha}}-\frac{d}{d t} \frac{\partial \psi^{v}}{\partial \dot{q}^{\alpha}}, \quad \dot{q}^{v}=\psi^{v}\left(q^{\alpha}, \dot{q}^{\alpha}\right) \tag{2.4}
\end{align*}
$$

$$
\left(\dot{q}^{3}=\dot{x}_{1}=\xi \dot{\varphi} \sin \varphi=\psi^{3}, \quad \dot{q}^{4}=\dot{y}_{1}=-\xi \dot{\varphi} \cos \varphi=\psi^{4}\right)
$$

Using (2.1)-(2.4), we obtain the following equations of motion

$$
\ddot{\xi}=0, \quad(\xi \dot{\varphi})^{\cdot}=0
$$

These equations can be integrated in succession giving

$$
\begin{align*}
& \dot{\xi}=C_{1}=\dot{\xi}_{0}, \quad \xi \dot{\varphi}=C_{2}=\xi_{0} \dot{\varphi}_{0}  \tag{2.5}\\
& \xi=\dot{\xi}_{0} t+\xi_{0}, \quad \varphi=\frac{\xi_{0} \dot{\varphi}_{0}}{\dot{\xi}_{0}} \ln \left(\dot{\xi}_{0} t+\xi_{0}\right)+C_{3} \tag{2.6}
\end{align*}
$$

In the special case when $\dot{\xi}_{0}=0$, the point $M_{2}$ will be at rest at the centre of a circle of radius $\xi_{0}$, while the point $M_{1}$ will move uniformly over this circle. On the other hand, if the point $M_{1}$ is at rest, the point $M_{2}$ will move uniformly along a fixed straight line, connecting these points.

We will put $\xi_{0}=1, \dot{\xi}_{0}=1, \varphi_{0}=0, \dot{\varphi}_{0}=1$

From Eqs. (2.2) and (2.6) we obtain the laws of motion of the points $M_{1}\left(s=1, \delta_{21}=0\right) 7$ and $M_{2}\left(s=2, \delta_{22}=1\right)$

$$
\begin{aligned}
& x_{s}=\frac{e^{t+1}}{2}\{\sin [\ln (t+1)]-\cos [\ln (t+1)]\}+\delta_{2 s}(t+1) \cos [\ln (t+1)] \\
& y_{s}=-\frac{e^{t+1}}{2}\{\sin [\ln (t+1)]+\cos [\ln (t+1)]\}+\delta_{2 s}(t+1) \sin [\ln (t+1)]
\end{aligned}
$$

## 3. Integrals that are linear in the momenta (cyclic coordinates)

We will show that, in the system considered, for each of the two linear integrals (2.5), we can choose generalized coordinates such that one of them will be cyclical (in the sense given earlier ${ }^{4}$ ), and the integral corresponding to it will be cyclical.

We will choose as the generalized coordinates of the system

$$
x_{1}=q^{1}, \quad y_{1}=q^{2}, \quad x_{2}=q^{3}, \quad y_{2}=q^{4}
$$

and we will write the constraint equations in the form

$$
f^{1}=\dot{q}^{1} \cos \varphi+\dot{q}^{2} \sin \varphi=0, \quad f^{2}=\dot{q}^{1} \dot{q}^{3}+\dot{q}^{2} \dot{q}^{4}=0
$$

The equations of motion allow of two first integrals that are linear and homogeneous in the velocities

$$
\begin{aligned}
& \left(q^{3}-q^{1}\right)\left(\dot{q}^{4}-\dot{q}^{2}\right)+\left(q^{4}-q^{2}\right)\left(\dot{q}^{3}-\dot{q}^{1}\right)=\text { const, } \\
& \frac{\left(q^{3}-q^{1}\right)\left(\dot{q}^{3}-\dot{q}^{1}\right)+\left(q^{4}-q^{2}\right)\left(\dot{q}^{4}-\dot{q}^{2}\right)}{\sqrt{\left(q^{3}-q^{1}\right)^{2}+\left(q^{4}-q^{2}\right)^{2}}}=\text { const }
\end{aligned}
$$

The Lagrange and Hamilton functions have the form

$$
L=\frac{1}{2}\left[\left(\dot{q}^{1}\right)^{2}+\left(\dot{q}^{2}\right)^{2}\right]+\frac{1}{2}\left[\left(\dot{q}^{3}\right)^{2}+\left(\dot{q}^{4}\right)^{2}\right], \quad H=\frac{1}{2}\left(p_{1}^{2}+p_{2}^{2}+p_{3}^{2}+p_{4}^{2}\right)
$$

We will choose as the new coordinates $Q^{1}$ and $Q^{2}$ the Cartesian coordinates of the point $M_{2}$, as $Q^{3}$ we will choose the length of the section $M_{1} M_{2}$ and we will choose as $Q^{4}$ the angle of inclination of the section $M_{1} M_{2}$ to the $x$ axis. The corresponding conversion formulae have the form

$$
q_{1}=Q^{1}-Q^{3} \cos Q^{4}, \quad q_{2}=Q^{2}-Q_{3} \sin Q^{4}, \quad q_{3}=Q^{1}, \quad q_{4}=Q^{2}
$$

The Lagrange and Hamilton functions, written, taking into account the constraint equations,

$$
\begin{aligned}
& F^{3}=\dot{Q}^{1} \cos Q^{4}+\dot{Q}^{2} \sin Q^{4}-\dot{Q}^{3}=0 \\
& F^{4}=\dot{Q}^{1} \sin Q^{4}-\dot{Q}^{2} \cos Q^{4}=0
\end{aligned}
$$

take the following form in the new variables

$$
L^{*}=\frac{1}{2}\left[\left(\dot{Q}^{1}\right)^{2}+\left(\dot{Q}^{2}\right)^{2}\right]+\frac{1}{2}\left(\dot{Q}^{3} \dot{Q}^{4}\right)^{2}, \quad H^{*}=\frac{1}{2} P_{1}^{2}+\frac{1}{2} P_{2}^{2}+\frac{1}{2} \frac{1}{\left(Q^{3}\right)^{2}} P_{4}^{2}
$$

Since

$$
\frac{\partial H^{*}}{\partial Q^{4}}=0, \quad A_{4}^{3}=\frac{\partial f^{3}}{\partial \dot{q}^{i}} \frac{\partial q^{i}}{\partial Q^{4}}=\frac{\partial F^{3}}{\partial \dot{Q}^{4}}=0, \quad A_{4}^{4}=\frac{\partial f^{4}}{\partial \dot{q}^{i}} \frac{\partial q^{i}}{\partial Q^{4}}=\frac{\partial F^{4}}{\partial \dot{Q}^{4}}=0
$$

the coordinate $Q^{4}$ is cyclical. We have the cyclic integral $P_{4}=\left(Q^{3}\right)^{2} \dot{Q}^{4}=$ const, i.e. $\xi^{2} \dot{\varphi}=$ const.
Choosing as the new coordinates $Q^{1}$ and $Q^{2}$ the Cartesian coordinates of the point $M_{1}$ and the coordinates $Q^{3}$ and $Q^{4}$, as shown above, we obtain the following conversion formulae

$$
q_{1}=Q^{1}, \quad q_{2}=Q^{2}, \quad q_{3}=Q^{1}+Q^{3} \cos Q^{4}, \quad q_{4}=Q^{2}+Q^{3} \sin Q^{4}
$$

As before, we conclude that the coordinate $Q^{3}$ is cyclical. We have the cyclic integral $P_{4}=\left(\dot{Q}^{3}\right)=$ const, i.e., $\dot{\xi}=$ const.

## 4. The stationarity of the Hamilton action

For the system being considered, we will check that the following conditions ${ }^{5}$ are satisfied

$$
\begin{equation*}
\lambda_{v} \gamma_{\alpha}^{v}=0 \tag{4.1}
\end{equation*}
$$

For these conditions the Hamilton principle will be the principle of stationary action. Here $\gamma_{\alpha}^{\nu}$ are terms of non-holonomicity in Eqs. (2.4) and $\lambda_{\nu}$ are coefficients of the Euler-Lagrange variational problem

$$
\delta \int_{t_{0}}^{t}\left(L+\lambda_{v} f^{v}\right) d t=0 ; \quad f^{v}=\dot{q}^{v}-\psi^{v}\left(q^{i}, \dot{q}^{\alpha}\right)=0
$$

The equations of the extremals

$$
\frac{d}{d t} \frac{\partial L}{\partial \dot{q}^{i}}-\frac{\partial L}{\partial q^{i}}=\lambda_{v}\left(\frac{\partial L}{\partial q^{i}}-\frac{d}{d t} \frac{\partial L}{\partial \dot{q}^{i}}\right)-\dot{\lambda}_{v} \frac{\partial f^{v}}{\partial \dot{q}^{i}}
$$

in general are not identical with the equations of motion of the system. They can be written in the form ${ }^{3}$

$$
\begin{align*}
& \frac{d}{d t} \frac{\partial L^{*}}{\partial \dot{q}^{\alpha}}-\frac{\partial L^{*}}{\partial q^{\alpha}}-\frac{\partial L^{*}}{\partial q^{v}} \frac{\partial \psi^{v}}{\partial \dot{q}^{\alpha}}+\left(\frac{\partial L}{\partial \dot{q}^{v}}+\lambda_{v}\right) \gamma_{\alpha}^{v}=0  \tag{4.2}\\
& \frac{d \lambda_{v}}{d t}+\frac{\partial \psi^{p}}{\partial q^{v}} \lambda_{p}+\frac{d}{d t} \frac{\partial L}{\partial \dot{q}^{v}}-\frac{\partial L}{\partial q^{v}}=0 \tag{4.3}
\end{align*}
$$

where

$$
L^{*}=L^{*}\left(q^{i}, \dot{q}^{\alpha}\right)=L_{\dot{q}^{\vee}=\psi^{v}}, \quad \gamma_{\alpha}^{v}=\frac{\partial \psi^{v}}{\partial q_{\alpha}}+\frac{\partial \psi^{\nu}}{\partial q^{p}} \frac{\partial \psi^{p}}{\partial \dot{q}^{\alpha}}-\frac{d}{d t} \frac{\partial \psi^{\nu}}{\dot{q}^{\alpha}}
$$

When conditions (4.1) are satisfied, Eqs. (4.2) take the form of the Voronets equations of motion of a non-holonomic system. Putting

$$
q^{1}=\xi, \quad q^{2}=\varphi, \quad q^{3}=x_{2}, \quad q^{4}=y_{2}
$$

we can represent the constraint equations in the form

$$
\dot{q}^{3}=\dot{\xi} \cos \varphi=\psi^{3}, \quad \dot{q}^{4}=\dot{\xi} \sin \varphi=\psi^{4}
$$

We will write the non-holonomicity coefficients (4.4) as follows:

$$
\begin{equation*}
\gamma_{1}^{3}=\dot{\varphi} \sin \varphi, \quad \gamma_{1}^{4}=-\dot{\varphi} \cos \varphi, \quad \gamma_{2}^{3}=-\dot{\xi} \sin \varphi, \quad \gamma_{2}^{4}=\dot{\xi} \cos \varphi \tag{4.4}
\end{equation*}
$$

Consequently, conditions (4.1) of the stationarity of the action are

$$
\begin{align*}
& \lambda_{3} \gamma_{1}^{3}+\lambda_{4} \gamma_{1}^{4}=\left(\lambda_{3} \sin \varphi-\lambda_{4} \cos \varphi\right) \dot{\varphi}=0 \\
& \lambda_{3} \gamma_{2}^{3}+\lambda_{4} \gamma_{2}^{4}=\left(-\lambda_{3} \sin \varphi+\lambda_{4} \cos \varphi\right) \dot{\xi}=0 \tag{4.5}
\end{align*}
$$

Dropping the trivial case $\dot{\xi}=\dot{\varphi}=0$, we reduce conditions (4.5) to the form

$$
\begin{equation*}
\lambda_{3} \sin \varphi-\lambda_{4} \cos \varphi=0 \tag{4.6}
\end{equation*}
$$

It follows from Eq. (4.3) that

$$
\lambda_{3}=C_{1}-\dot{x}_{2}, \quad \lambda_{4}=C_{2}-\dot{y}_{2} \quad\left(C_{1}, C_{2}-\text { are constants }\right)
$$

and hence limitation (4.6) (taking the constraints into account) reduces to the equation

$$
C_{1} \sin \varphi-C_{2} \cos \varphi=0
$$

i.e., $\varphi=$ const. So, the particular motion in which the point $M_{1}$ is at rest while the point $M_{2}$ moves along a straight line, yields a stationary value of the Hamilton action.If we take as the generalized coordinates

$$
q^{1}=\xi, \quad q^{2}=\varphi, \quad q^{3}=x_{1}, \quad q^{4}=y_{1}
$$

the constraint equations can be written as

$$
\begin{equation*}
\dot{q}^{3}=\xi \dot{\varphi} \sin \varphi=\psi^{3}, \quad \dot{q}^{4}=-\xi \dot{\varphi} \cos \varphi=\psi^{4} \tag{4.7}
\end{equation*}
$$

where the terms of non-holonomicity retain the form (4.4), and, consequently, the stationarity conditions (4.5) do not change. We obtain from Eqs. (4.3)

$$
\lambda_{3}=C_{0}-\dot{x}_{1}, \quad \lambda_{4}=C_{1}-\dot{y}_{1}
$$

as a consequence of which condition (4.6) becomes the equation

$$
\lambda_{3} \sin \varphi-\lambda_{4} \cos \varphi=C_{0} \sin \varphi-C_{1} \cos \varphi-\xi \dot{\varphi}=C_{0} \sin \varphi-C_{1} \cos \varphi-C_{2}=0
$$

(the second integral of (2.5) has been used). It is only satisfied for the already known case $\varphi=$ const.
However, there is one more case when, in the particular motion of the system, the Hamilton action takes a stationary value. This case is $\dot{\xi}=0$ : the point $M_{2}$ is at rest while the point $M_{1}$ moves in a circle. The point is that in this case the constraints (4.7) are integrable, and the system is holonomic, as a result of which the action stationarity conditions are satisfied. This case is not obtained from the preceding analysis due to the structure of the coefficients $\gamma_{\alpha}^{\nu}$ (4.4), because the coefficient $\xi$ occurs in terms the sum of which is equal to zero.

## References

1. Zeković D. Examples of non-linear non-holonomic constraints in classical mechanics. Vestnik MGU. Ser. 1. Matematika Mekhanika 1991;1:100-3.
2. Neimark YuI, Fufayev NA. The Dynamics of Non-holonomic Systems. Moscow: Nauka; 1967.
3. Zeković D. Some problems of the dynamics of non-holonomic systems. Doctorate Dissertation, Beograd; 1984.
4. Zeković DN. Linear integrals of non-holonomic systems with non-linear constraints. Prikl Mat Mekh 2005;69(6):929-34.
5. Rumyantsev VV. Hamilton's principle for non-holonomic systems. Prikl Mat Mekh 1978;42(3):387-99.

[^0]:    [3 Prikl. Mat. Mekh. Vol. 72, No. 5, pp. 721-726, 2008.

